

Lecture 30 Bose Einstein Condensation 4/20/01

See [jila-www.colorado.edu/bec](http://jila-www.colorado.edu/bec)

First formed by Carl Weiman's group at Univ. of Colorado in 1995  
Predicted by Bose + Einstein in 1924

Critical density  $\Rightarrow$  density of Bose gas as  $\mu \rightarrow 0$

$$\rho_{\text{crit}} \sim \frac{1}{\lambda^3}$$

$\lambda =$  thermal deBroglie wavelength  $= \frac{h}{p}$

$$\frac{p^2}{2m} \sim k_B T$$

$$\rho_{\text{crit}} \sim \left( \frac{2m k_B T}{h^2} \right)^{3/2}$$

- Need either very high density or very low temperature because  $h$  is very small.

Assume ideal Bose gas (neglect interactions between atoms)

$$P_{\text{crit}} = g \int \frac{d^3 k}{(2\pi)^3} \left[ \frac{1}{e^{E/k_B T} - 1} \right]$$

$$g = \text{spin degeneracy} = 2s + 1$$

$$= 1, 3$$

$$s = 0, 1$$

Bose gas  $s$  must be integer

$$E = \frac{\hbar^2 k^2}{2m}$$

$$\text{let } x = \frac{\hbar^2 k^2}{2m k_B T}$$

$$dx = \frac{\hbar^2}{2m k_B T} 2k dk$$

$$P_{\text{crit}} = \frac{4\pi}{(2\pi)^3} g \left( \frac{2m k_B T}{\hbar^2} \right)^{3/2} \frac{1}{2} \int_0^\infty \frac{x^{1/2} dx}{e^x - 1}$$

$$k = \left( \frac{2m k_B T}{\hbar^2} \right)^{1/2} x^{1/2}$$

$$P_{\text{crit}} = \frac{4\pi g}{(2\pi)^3} \int_0^\infty \frac{k^2 dk}{e^{E_k/k_B T} - 1}$$

$$= \frac{g}{2\pi^2} \left( \frac{2m k_B T}{\hbar^2} \right)^{3/2} \frac{1}{2} \int_0^\infty \frac{x^{1/2} dx}{e^x - 1}$$

$$\int_0^\infty \frac{x^{s-1} dx}{e^x - 1} = \Gamma(s) \zeta(s)$$

[See 5.109]

$$\zeta\left(\frac{3}{2}\right) = \frac{\pi^{1/2}}{2}$$

$$\zeta(s) = \sum_{p=1}^{\infty} p^{-s}$$

$$\zeta(3/2) = 2.612 \dots$$

$$P_{\text{crit}} = \left[ \frac{g \zeta(3/2)}{8 \pi^{3/2}} \right] \left( \frac{2 m k_B T}{h^2} \right)^{3/2}$$

Example Liquid  ${}^4\text{He}$  at  $T = 2 \text{ K}$

${}^4\text{He}$  has  $2p + 2n + 2e^-$

Can add an even # of spin  $\frac{1}{2}$  particles to get an integer spin  
 $\Rightarrow {}^4\text{He}$  is a boson.

It has  $S = 0 \Rightarrow g = 1$

$$M \approx 4 (1.661 \times 10^{-27} \text{ kg}) = 6.64 \times 10^{-27} \text{ kg}$$

$$k_B = 1.381 \times 10^{-23} \text{ J/K}$$

$$h = 1.055 \times 10^{-34} \text{ J}\cdot\text{s}$$

$$P_{\text{crit}} = \left( \frac{2.612}{8 \pi^{3/2}} \right) \left[ \frac{2 \cdot 6.64 \times 10^{-27} \cdot 1.38 \times 10^{-23} \cdot 2}{(1.055 \times 10^{-34})^2} \right]^{3/2}$$

$$= .0586 (1.89 \times 29) = 1.107 \times 28$$

$$\begin{aligned}
 \rho_{\text{crit}} &= 1.11 \times 10^{28} \text{ particles / m}^3 \\
 &= 1.11 \times 10^{22} \frac{\text{particles}}{\text{cm}^3} \\
 &= 1.11 \times 10^{22} \cdot 6.64 \times 10^{-24} \text{ g} = 0.074 \text{ gm/cm}^3
 \end{aligned}$$

This is close to the real density of superfluid  $^4\text{He}$   $\rho \approx 0.15 \text{ gm/cm}^3$

Note  $^4\text{He}$  is not a pure Bose condensate because interactions between  $^4\text{He}$  atoms are very strong.

However transition to superfluid phase ~~is~~ at  $T = 2.17 \text{ K}$  is associated with the formation of a zero momentum condensate.

IF you cool superfluid  $^4\text{He}$  to very low temperatures only about 10% of the atoms go into the condensate

To make a real Bose-Einstein condensate cool a very dilute gas to very low temperatures

If  $\lambda \gg r_i$

where  $r_i$  is the range of the interactions between atoms than system will behave as an ideal Bose gas.

In 1995  $^{87}\text{Rb}$  atoms were cooled in a magnetic trap.

Used several methods of cooling

### Laser cooling

Use a laser with frequency just below an atomic transition



atom

at rest

or moving

away does not absorb light

But an atom moving towards laser sees doppler shifted light beam and absorbs light, Momentum of light

photon slows atom down.

$$p = \hbar \omega / c$$

Cooled system to  $\sim 170$  nanoK

$$T = 170 \times 10^{-9} \text{ K}$$

$^{87}\text{Rb}$  atoms have an odd # of electrons plus an odd # of neutrons + protons  $\rightarrow$  sum to an integer spin (boson)  
 $Z = 37$   $e^-$  + 37 protons + 50 neutrons

$$M \approx 87 (1.66 \times 10^{-27}) \text{ kg} = 1.44 \times 10^{-25} \text{ kg}$$
$$k_B T = 1.38 \times 10^{-23} \text{ J/K} \quad 170 \times 10^{-9} = 2.35 \times 10^{-30} \text{ J}$$
$$\hbar = 1.05 \times 10^{-34} \text{ J/s}$$

$$\rho_{\text{crit}} = 0.0586 \left[ \frac{2 \cdot 1.44 \times 10^{-25} \cdot 2.35 \times 10^{-30}}{(1.05 \times 10^{-34})^2} \right]^{3/2}$$
$$= 2.8 \times 10^{19} \text{ particles / m}^3$$

$$\rho_{\text{crit}} = 2.8 \times 10^{13} \text{ atoms / cm}^3$$

Colorado group claimed slightly lower critical density but not on pillar. ~~Am. Phys. Monthly~~